Handout 11: p-modes for the isentropic case

1 Derivation for isentropic case

The continuity equations is $D \ln \rho / Dt = -\nabla \cdot \boldsymbol{u}$, and from Ds / Dt = 0 we obtain $D \ln p / Dt = -\gamma \nabla \cdot \boldsymbol{u}$. Our linearized equations are then

$$\dot{\rho_1} + \boldsymbol{u}_1 \cdot \boldsymbol{\nabla} \rho_0 + \rho_0 \boldsymbol{\nabla} \cdot \boldsymbol{u}_1 = 0, \tag{1}$$

$$\dot{p}_1 + \boldsymbol{u}_1 \cdot \boldsymbol{\nabla} p_0 + \gamma p_0 \boldsymbol{\nabla} \cdot \boldsymbol{u}_1 = 0, \tag{2}$$

$$\rho_0 \dot{\boldsymbol{u}}_1 + \boldsymbol{\nabla} p_1 + \rho_1 g \hat{\boldsymbol{z}} = 0. \tag{3}$$

Assume $\dot{\boldsymbol{\xi}} = \boldsymbol{u}$ and integrate the continuity equation in time, i.e.,

$$\rho_1 + \boldsymbol{\xi}_1 \cdot \boldsymbol{\nabla} \rho_0 + \rho_0 \boldsymbol{\nabla} \cdot \boldsymbol{\xi}_1 = 0, \tag{4}$$

$$p_1 + \boldsymbol{\xi}_1 \cdot \boldsymbol{\nabla} p_0 + \gamma p_0 \boldsymbol{\nabla} \cdot \boldsymbol{\xi}_1 = 0, \tag{5}$$

$$\rho_0 \hat{\boldsymbol{\xi}}_1 + \boldsymbol{\nabla} p_1 + \rho_1 g \hat{\boldsymbol{z}} = 0. \tag{6}$$

Divide by ρ_0 , use $\nabla p_0 = -\rho_0 g \hat{z}$, $\gamma p_0 = c_s^2 \rho_0$, replace $g \hat{z} = c_s^2 \hat{z} / H_\rho = -c_s^2 \nabla \ln \rho_0$, and insert the above expressions for p_1 and ρ_1 , so

$$\rho_1 + \boldsymbol{\xi}_1 \cdot \boldsymbol{\nabla} \rho_0 + \rho_0 \boldsymbol{\nabla} \cdot \boldsymbol{\xi}_1 = 0, \tag{7}$$

$$p_1 - \xi_{1z} \rho_0 g + c_s^2 \rho_0 \nabla \cdot \boldsymbol{\xi}_1 = 0,$$
(8)

$$\ddot{\boldsymbol{\xi}}_{1} + \frac{1}{\rho_{0}} \boldsymbol{\nabla} p_{1} + \frac{\rho_{1}}{\rho_{0}} g \hat{\boldsymbol{z}} = 0.$$
(9)

Insert p_1 and ρ_1 into the momentum equation, so

$$\ddot{\boldsymbol{\xi}}_{1} + \frac{1}{\rho_{0}} \boldsymbol{\nabla} \left[\rho_{0}(\boldsymbol{\xi}_{1z}g - c_{\mathrm{s}}^{2}\boldsymbol{\nabla} \cdot \boldsymbol{\xi}_{1}) \right] - (\boldsymbol{\xi}_{1} \cdot \boldsymbol{\nabla} \ln \rho_{0} + \boldsymbol{\nabla} \cdot \boldsymbol{\xi}_{1}) g \hat{\boldsymbol{z}} = 0, \tag{10}$$

or

$$\ddot{\boldsymbol{\xi}}_{1} + \frac{1}{\rho_{0}} \boldsymbol{\nabla} \left[\rho_{0} (\boldsymbol{\xi}_{1z}g - c_{\mathrm{s}}^{2} \boldsymbol{\nabla} \cdot \boldsymbol{\xi}_{1}) \right] + (\boldsymbol{\xi}_{1z}/H_{\rho} - \boldsymbol{\nabla} \cdot \boldsymbol{\xi}_{1}) g \hat{\boldsymbol{z}} = 0, \tag{11}$$

or, using the fact that $gH_{\rho} = c_{\rm s}^2$, we get

$$\ddot{\boldsymbol{\xi}}_{1} + \frac{1}{\rho_{0}} \boldsymbol{\nabla} \left[\rho_{0} (\boldsymbol{\xi}_{1z}g - c_{\mathrm{s}}^{2} \boldsymbol{\nabla} \cdot \boldsymbol{\xi}_{1}) \right] + (\boldsymbol{\xi}_{1z}g - c_{\mathrm{s}}^{2} \boldsymbol{\nabla} \cdot \boldsymbol{\xi}_{1}) \hat{\boldsymbol{z}} / H_{\rho} = 0,$$
(12)

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$$\ddot{\boldsymbol{\xi}}_1 + \boldsymbol{\nabla} \left(\xi_{1z}g - c_{\rm s}^2 \boldsymbol{\nabla} \cdot \boldsymbol{\xi}_1 \right) = 0.$$
(13)

Assuming that the displacement is a potential field, i.e., $\boldsymbol{\xi}_1 = \boldsymbol{\nabla} \Phi$, we have (Bogdan & Cally, 1995)

$$\ddot{\Phi} + g\partial\Phi/\partial z - c_{\rm s}^2 \nabla^2 \Phi = 0.$$
⁽¹⁴⁾

This leads to an eigenvalue problem for eigenvalue ω^2 ,

$$\omega^2 \Phi = c_{\rm s}^2 k^2 \Phi + g \frac{\partial \Phi}{\partial z} - c_{\rm s}^2 \frac{\partial^2 \Phi}{\partial^2 z}.$$
(15)

At the bottom, we assume $\Phi = 0$, which means that the last data point has to be omitted from the matrix. At the top, we require $(k^2 - d^2/dz^2)\Phi = 0$, which means that we replace the eigenvalue problem on the boundary by

$$\omega^2 \Phi = g \frac{\partial \Phi}{\partial z} \tag{16}$$

Note that Equation (15) is valid even in the non-isothermal case. In Figure 1, we compare solutions for the isothermal and non-isothermal (n = 3/2) cases. Note that in the latter case the *p*-modes bend down and the difference between qualitative difference between the *f*- and *p*-modes diminishes.

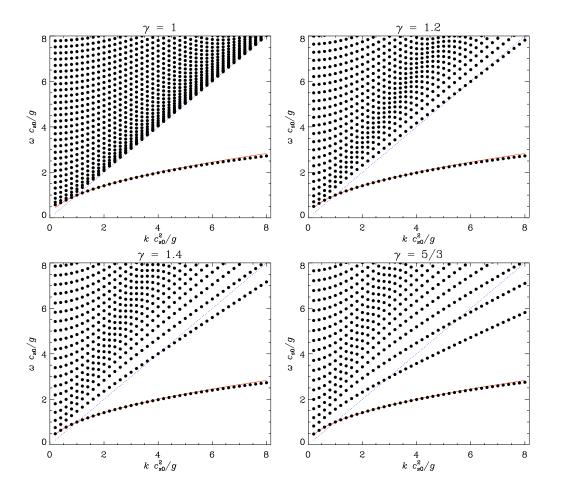


Figure 1: $k\omega$ diagram for isothermal case (left) and various isentropic cases.

2 Reduction to the isothermal case

It is instructive to see how to obtain the dispersion relation for the isothermal case from Equation (15). Replacing $\partial_z \to ik_z$, we have

$$\omega^2 = c_{\rm s}^2 k^2 + ik_z g + c_{\rm s}^2 k_z^2 = 0, \qquad (17)$$

Split $k_z = k_z' + ik_z''$ into real and imaginary parts, so we get

$$\omega^{2} = c_{\rm s}^{2}k^{2} + (\mathrm{i}k_{z}' - k_{z}'')g + c_{\rm s}^{2}(k_{z}'^{2} - k_{z}''^{2} + 2\mathrm{i}k_{z}'2k_{z}'') = 0,$$
(18)

and solve for real and imaginary parts:

$$\omega^2 = c_{\rm s}^2 k^2 - k_z'' g + c_{\rm s}^2 (k_z'^2 - k_z''^2) = 0, \qquad (19)$$

$$gk'_z + 2k'_z 2k''_z c_s^2 = 0, (20)$$

So now replace $k_z'' = -g/2c_s^2$ into Equation (19) and get

$$\omega^2 = c_{\rm s}^2 k^2 + g^2 / 2c_{\rm s}^2 + c_{\rm s}^2 (k_z'^2 - g^2 / 4c_{\rm s}^4) = 0, \qquad (21)$$

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$$\omega^2 = c_{\rm s}^2 k^2 + g^2 / 4c_{\rm s}^2 + c_{\rm s}^2 k_z'^2 = 0.$$
(22)

References

Bogdan, T. J., & Cally, P. S., "Jacket Modes: Solar Acoustic Oscillations Confined to Regions Surrounding Sunspots and Plage," Astrophys. J. 453, 919-928 (1995).

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